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No. XIX.

A DISSERTATION

ON THE

ORIGIN OF SPRINGS.

By SERENO DWIGHT, Esq.

PHILOSOPHY has long been defirous of investigating the Causes of things; but has usually made slow progress, where the modus operands, or the connecting link between cause and effect, could neither be seen nor felt. To explain the phenomena of Thunder and Lightning, she first created a Jupiter, and then forged his thunderbolts. For the cause of Tempests, she resorted to the influence of the Stars. And Herodotus very gravely tells us, that the return of the Sun from the south, after the winter solftice, is owing to the prevalence of a violent South Wind in Egypt.

The Origin of Springs is a subject of this invisible and impalpable nature; and, as might have been expected, has long agitated the Republic of Letters. ARISTOTLE informs us, that the air, which is inclosed in the vaults and caverns of the Earth, is condensed into water near the surface; and thence flows out in springs. In the present state of science, probably, no one will think, that this scheme needs a resutation. Many au-

thors, however, who have treated it with an unbecoming ridicule, would have fpared their wit, had they known, that their own hypotheses, devised more than two thousand years afterwards, were destined to a similar fate with that of the Grecian Philosopher.

The Modern Theories may be reduced to three classes.

They all find the refources of Springs

In the Ocean;

In an Abys, in the bowels of the Earth; or

In Vapour.

Those, who draw the water of Springs from the Ocean, have been puzzled to convey it to their orisices in the state in which we actually find it. Springs are usually fresh: the Ocean is falt. Most Springs are higher, and some many thousand feet higher, than the level of the Ocean. To raise their supplies requires, therefore, a force sufficient to counteract the force of

Gravity.

These difficulties, attending the first class of theories, have cost its advocates no small expence of labour and ingenuity. Des Cartes, to avoid them, kindled a fire in the bowels of the Earth, by which he converted the water of the Ocean into Vapour. This Vapour, he collected in hollow subterranean caverns, and these condensed it into water. Some of his followers, dispensing with these caverns, raised their vapour thus formed through the interstices of the ground, until it was condensed by cold near the surface. Had Des Cartes and his disciples adopted the rule, which afterwards directed the researches of Newton;

Never to afcribe a phenomenon to any cause, until the existence of that cause is proved; their subterranean fire would never have been kindled. Perhaps, also, they would have found the remote cause of springs in the heat of a very different fire; kindled, not merely in the imagination,—and by the hand of One, all whose

theories are only practical.

The existence of this subterranean fire is a mere hypothesis,—wholly unsupported by proof. So far as the

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interior of the Earth has been explored, its temperature below 1000 feet has been found, in all latitudes, to be about 50° of Fahrenheit: a temperature, which will appear wholly inconfiftent with the existence of fuch a fire, if we attend to the following facts.— The Rivers now running have probably flowed ever fince the Deluge; and the quantity of water, which they now discharge, has doubtless been their average fupply. Vapour is found to lofe 1000° of heat, when condenfed to water: and water absorbs the same quantity, when reconverted into vapour. According to the theory, the vapour must rife, before it is condensed, at least as high as the Springs which it supplies. What an incalculable quantity of heat must then have been discharged near the surface of the Earth, in order to fupply the rivers with water fince they first began to flow. This heat must have been constantly accumulating near the furface, during fo long a period. Why, then, is it no where discovered?

Others have attempted to avoid these objections, by calling in the aid of capillary attraction. As water is known to ascend in glass tubes of a very small bore, it is concluded that it may also ascend through ducts in the Earth of a similar size. Hence the existence of such ducts, and the ascent of water through them, are taken for granted. The Springs, which feed the Amazon, are several thousand miles from the Ocean; and many of them issue from two to four miles above its level. He, who can believe, that the interior of the Earth is thus surnished with an apparatus of natural capillary aqueducts running horizontally thousands of miles, till they meet a mountain sit to harbour a spring, and then ascending perpendicularly to an orisice; has

faith enough and to spare.

But if these ducts actually existed, they would not explain the phenomena of Springs.—To whatever cause the rise of water in capillary tubes is owing; whether to the attraction of the whole interior surface of the tube; or to the attraction of the ring of glass contiguous to the

upper furface; Reason demonstrates, that such a tube cannot raise more water than it can sustain; fince the nearer the fluid is, the stronger is the attractive force. Experiment confirms this conclusion; for a capillary tube, however short, cannot be made to run over.-The force of capillary attraction, also, will raise water but to a very moderate height. The proposition found in most Treatifes on Hydrostatics-That the height to which fluids rife in capillary tubes, is inverfely as the diameter of their bores—has led many to the erroneous conclusion, that if the bore is indefinitely small, the fluid will afcend to an indefinite height. But the fact is otherwife. Dr. Hooke, after many experiments with what he calls cobweb tubes, was unable to raife water in them more than 21 inches.—Should we, therefore, suppose such ducts to exist in the earth of the requisite length and position, capillary attraction would not raise water in any of them to a greater height than 21 inches; nor, even in those of a less length, would it cause the fluid to run over.

Others allege, that if a small heap of ashes or fand is put into a bason of water, the water will-rise through it, above its own level, to the top of the heap. They compare the Dry Land to the heap of ashes; and the Ocean to the water of the bason; and insist that the water of Springs rifes in the same manner through hills and mountains.-I am not informed, that any experiments have been tried to afcertain the height, to which. water will thus afcend through ashes, or through fand. The cause of its ascent, however, is well known to be a mutual attraction subfifting between the particles of ashes or fand, and those of water. Whether this attraction be chemical or mechanical, it is to me felf evident, that no attraction can raise more water than it can fustain. - If it is chemical, it will continue to raise the fluid until the heap is thoroughly faturated, and no longer. Should the other particles of the fluid attempt to supplant those which were first combined, these would effectually refult the attempt by the right of occupancy. If the attraction is merely mechanical, its force is well known to diminish, as the square of the distance increases. The heap, therefore, being supposed to have raised as much water as it can sustain; it is obvious, that the particles of water actually raised will be in immediate contact with the ashes or sand; while all the other particles will be at a greater distance from it. The water raised will thus be more strongly attracted, than the water not raised. Of course, the ashes or sand will continue to sustain the particles sirst raised: and will not be compelled, by a weaker force, to let them go, and thus make room for others. In other words, a feebler attraction can not overcome one that is more

powerful.

But it is faid, that if we conceive of angular pipes or ducts in the Earth, having the form, polition, and properties, of a syphon; the water of the Ocean may ascend and be discharged through them, by the same laws which regulate that instrument. We will admit, for the argument, that it is strictly philosophical to conceive of ducts or pipes in the Earth bermetically tight, and of a fufficient extent to answer the purpose. This hypothesis will, nevertheless, be attended with an infuperable difficulty. It is well known, that the difcharging orifice of the fyphon must be horizontally lower than the furface of the refervoir. In the case suppofed the Ocean is the refervoir. Of course, no spring, that is not fomewhat below the level of the Ocean, can be fupplied by a natural fyphon. With the few fprings fo fituated, we will not embarrafs our enquiries : fatiffied, that that wisdom, which operates by general laws; and, by the simplest means, produces the most magnificent refults; when it had filled the " Upper Springs, could find no difficulty in supplying the "nether."

The Wet Rag, like the Syphon, has been called in to relieve the perplexities of Philosophy. It is well known, that if such a rag is thrown partly over the side of a vessel of water, the water will drop from the exterior end of the rag, until the whole is thus drawn off. The

Wet Rag in this case is a sort of clumfy natural syphon. But, like the more regular instruments of art, it is faithfully true to its principles of action. Unless the surface of the fluid is above the horizontal level of the exterior end of the Rag, it is found, that not a drop of water will fall.

None of these various schemes, therefore, if they would explain the ascent of water to the orifices of

fprings, will account for its running over.

Nor will either of them furnish the reason, why the water of springs is fresh. The water, when it begins to ascend, is mere brine; and is supposed to lose its falt in the passage. If it were true, that brine would thus become fresh, it would follow, that the lowest stratum of earth, or that nearest the Ocean, would be immediately, and the fuperior strata gradually, faturated with falt; and both would of course decline receiving any more from the brine, as it passed them on its way upward to the fountains. The brine would thus freshen less and less in its passage; and in the end would continue wholly falt. The lower strata, also, would foon be faturated with folid falt; that either the tubes, or the interftices, or the fyphons, would become completely clogged, and incapable of any farther transmission of brine.

But it is not true, that fea-water is made fresh, by filtering through dry sand or dry earth. Experiment has proved, again and again, that, if brine is filtered through dry earth any number of times, its saltness is not perceptibly diminished. The quantity of sluid is indeed lessened; but that, which remains at the end of the process, is found to retain its original proportion of salt; and that, which is retained by the sand, adheres to it in the shape of brine, and not of solid salt. These considerations have satisfied me, that we must look to some other source than the Ocean, for the water of Springs.

The fecond class of Theories comprises those, which attribute the Origin of Springs to a vast Abyss in the

bowels of the Earth.

When Philosophy discovered this Abyss, her motives were praise-worthy, whatever we may think of her Logick. Infidelity had often attacked the Scriptural account of the Deluge, on the ground, that all the water on the globe was not fufficient to cover its furface to the depth represented by Moses. This objection, if the fact it afferts were true, would rest on the unfounded principle, that the CREATOR of all things is dependent on the things that are made for the accomplishment of his purposes. Some well meaning friends of the Penteatuch, alarmed for the credit of Moses, devifed this Abyss as the receptacle, in which the waters of the Deluge were gathered, that they might no more overflow the face of the Earth. Several philosophers, who had been put to great difficulty to account for the Origin of Springs, finding, in their fubterranean refearches, fo copious a refervoir prepared to their hands; immediately feized upon it as the fource, whence they were fupplied with water.

The best account I have seen of this Abyss, and of the manner in which sountains are sed by it, is sound in Catcott's Treatise on the Deluge: a work declared by Jones and Adams, two distinguished philosophers of Great Britain, one of the last, the other of the present century, to be "the most critical and satisfactory discourse extant on the origin of Springs and Rivers."

Mr. CATCOTT explains his own view of the internal fructure of the Earth, as it has existed fince the Deluge, by an engraving, representing the plane of one of its Great Circles. "At the centre we find," to use his own language, "a folid ball, or Nucleus, of terrestrial matter, formed from what the water of the Deluge, in its descent from the surface and passage through the strata of the Earth, tore off, and carried down with it into the Abys, and reposited at the lowest place. Around this Nucleus is the great Abys of water, with which all seas, lakes and rivers communicate. This Abys contains so large a quantity of water, that only a small part of it was used at the Deluge. Lastly, we

find the Crust of the Earth surrounding the Abys, and broken into innumerable apertures and fissures; the largest of which are the beds of Lakes, Seas and Oceans; the next less are the canals for the waters of Springs; and the least, the cracks, through which the vapours of the Abys ascend into the Atmosphere." To render this account more intelligible the Author compares the Earth to an Egg: its Crust to the Shell; the Abys, to the White; and the Nucleus, to the Yolk. He then enters into an elaborate argument to prove the existence of this Abys of waters; an attempt, in which he is pronounced successful by Jones and Adams.

Three methods are pointed out by Mr. CATCOTT, in which the waters of the Abyss may be conveyed to the

orifices of Fountains:

By the afcent of Vapour through the Cracks in the Crust;

By Upward Filtration; and By the Pressure of the Ocean.

The Vapour, mentioned in the first method, is supposed to be occasioned by a fire, existing somewhere in the bowels of the Earth. To avoid repeating what has already been faid respecting an internal fire; I will barely remark, that fuch a fire can not exist in the Nucleus; for there can be no regular supply of air to support it: nor in the Abys; for, had it once been kindled there, the water must immediately have extinguished it: and that if supposed to exist in the Crust, it could have no effect on the waters beneath. Were we, however, to allow the existence of such a fire in the Nucleus, operating on the bottom of the Abyls, like our common culinary fires on the bottom of a kittle : ftill, the Vapour thus occasioned would be condensed, in its ascent through many miles of cold water, long before it reached the under furface of the Crust. The only alternative is to imagine the Abyss to be, and to have always been, an immense subterranean ocean of hot water. - In this way, if we suppose 1st. That there is fuch an abysis; 2dly. That it is a mere mass of hot water; 3dly. That the Crust of the Earth has the requisite number of cracks and siffures; and 4thly. That, near the surface, the vapour finds the proper caverns or refrigeratories to condense in: we shall have made but two more suppositions than the Indian, who put the Earth on an Elephant, and the Elephant on a Tortoise.

If Vapour is thus continually rifing from the Abyss, why do we never fee it making its way above the furface in a visible shape? To this question, Mr. CAT-COTT replies, " As to the afcent of Vapour through the fiffures of the Earth, this is a fact, the reality of which any one may be convinced of, who will give himself the trouble of looking into the infide of the Earth." Of the Earth, most of us, probably, are merely superficial observers. For myself, unfortunately, I never had an opportunity of examining its infide except while exploring a cave about 70 feet deep. In that, and in most other caverns and mines, water is feen dropping from the roof, and trickling down the fides, of those fubterranean recesses; and the atmosphere, thus continually moistened, has no chance of being dried by the heat of the Sun. Miners, however, and most other vilitors of caverns, guided by common fense, have been led to attribute the moisture to these causes, and not to fubterranean exhalations.

The fecond method devised for the ascent of the waters of the Abys is Upward Filtration. I flatter myfelf, that it has been already evinced—that water cannot, thus, ascend to a sufficient height:—That, if it could, it would not run over; and that, supposing it would run over, yet, if originally salt, it could not be

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With regard to the quality of the water of the A-byss, Mr. Catcott observes, that "we cannot have any absolute proof that it is falt; and I could give several reasons to show that it may not be so; at least, not equally falt with the sea." What these reasons are he does not inform us. Perhaps it is owing to this o-

mission, that several reasons of a contrary efficacy have fatisfied me, that the Abyss, even if it were originally fresh, must now consist of falt water. The Deluge prevailed on the Earth one hundred and fifty days. The waters of the Abyss, according to Mr. C. were employed, as subsidiary to those of the Ocean, in drowning the dry land; and then were returned to their appointed bed. Great indeed must have been the care, taken during that long period, to prevent the brine of the Ocean from intermixing with the fresh water of the Abyss.

The Ocean, also, according to Mr C. communicates directly with the waters of the Abyss; or in other words, is a part or continuation of it. If this be the case, the waters of the Abyss must be at least as falt as those of the Ocean; for falt water is heavier than fresh, and will fink in it until there is produced an equilibri-

um of fpecific gravity,

If the Ocean is not a part of the Abys, but refts upon a substratum of earth, which in its turn rests upon the surface of the Abys; then, no reason can be assigned why a due proportion of the cracks and sissures in the crust should not be assigned to the bed of the Ocean. Only one sist of the surface of the globe is dry land. The remaining four sists are covered by the Sea. The Dry Land, according to the hypothesis, has as many such sissures as it has Springs. And singular indeed must have been the care necessary after the Deluge, in settling the Wreck of Elements, to distribute these sissures in such a manner, that so many should be found in the Land, and not one in the bottom of the Ocean.

If these cracks and fissures are impartially distributed over the Crust of the Earth; it is clear, that the waters beneath must be as falt as those above. Let us suppose that, during the Deluge, the waters of the Abyss, by a strange coyness, were preserved from contamination; and that, after their bring ordeal, they returned pure to their native bed. Still they could only have returned to meet dangers insurmountable. The brine of

the Ocean, being specifically heavier, must immediately have begun to descend through some of the cracks, and the water of the Abyss to ascend through others. Nor could the process have ceased, until it had produced an equilibrium of pressure. If this rotation in the waters of our Great Globe is not yet over, it will account, in a manner which HALLEY never dreamed of,

for whirlpools, edies, and waterspouts.

The third mode contrived for the afcent of the water of the Abys is the pressure of the Ocean. This Pressure is thus explained.—A crack or sissure passes downward, from the orisice of every spring, through the Crust, to the surface of the Abys, which is supposed to be fresh. If water is poured into one arm of a bended tube, and oil into the other; the level of the oil, as being the lighter sluid, will remain higher than that of the water. The whole cavity of the Ocean may be considered as one arm of such a tube; a given crack or sissure, the other; and the Abys, the connection between them. The brine of the Ocean, pressing upon the lighter sluid of the Abys, will force it up through the sissure, to a greater height than its own level, and thus will form a spring.

On this scheme, the Ocean must be connected with the Abyss; for, if it is not, it cannot press upon it. If it is connected with it, then I think it has been proved, that the waters of the latter must be falt. If they are falt, the pressure of the Ocean cannot raise them above its own level; for, if water alone is poured into a bended tube, one arm of which is a hundred or a thousand times as large as the other, still the surface of the fluid in both will have the same horizontal level.

But if this preffure would raise the waters of the A. byss to a sufficient height; still, it would not make them fresh. The cracks or sissures are mere tubes of a palpable diameter; and to freshen brine, merely by passing it through a tube, is a harder problem than a discreet Chemist would attempt to solve.

The manner in which springs issue from the ground,

also, completely disproves the theory. Springs make their way nearly to the tops of the Andes. If the pressure of the Ocean is sufficient to force the water of the Abyss to such a height; but little of its force can have been spent in such springs as issue on plains. In these, therefore, the water, instead of merely gurgling upwards, as they actually do, should rise in jets; and contine rising, until the impulse communicated by this pressure was overcome by the resistance of the air, and

the Force of Gravity.

As to the length of the cracks or fiffures connecting Springs with the Abyss, it should be remarked that the Crust or Shell of the Earth cannot, on any reasonable fupposition, be considered as less than several miles in thickness. In a Globe of 8000 miles diameter, with a centrifugal force at its equator fo powerful as that of the Earth; it will be readily felt by every man, that unless the Crust were folid and substantial, there would be constant and imminent hazard of something worse than a mere fiffure. Springs, also, are often two or three miles above the level of the Ocean. Each Spring must have its own fiffure commencing at the furface of the Abyss, and reaching through the Crust to the orifice. The man, who can believe that the SUPREME BEing could devife no eafier and better mode of watering the Earth, must I think, be in great danger of thinking him, so far as his wisdom is concerned, altogether such an one as himfelf.

Leaving, then, our fubterranean refearches, we will examine the Theory, which discovers the Origin of

Springs in Vapour.

Vapour, as here used, is a comprehensive term, including all the water, that rises from the surface of the Earth by evaporation; and all, that descends on it in the form of dew, fog, mist, rain, snow, and hail.

This Theory may be thus explained. The Ocean is conftantly losing vast quantities of water by evaporation. Electrified bodies attract light substances floating near them. The Land is more highly electrified than the

Ocean. Most of the vapour thus raised is drawn from the fea to the land. As mountains are more highly e. lectrified than plains, they attract the great body of the vapour which retains the form of mist; while of that, which descends in a more folid form, much more falls on the former than on the latter. The mift on mountains condenses, and is precipitated in water. This united with the water of rain and fnow, penetrates the strata of fand and the lighter earth, till it is stopped in its course by more impervious substances; particularly strata of clay. In these, it forms a bason or reservoir: from which, gradually working a paffage, it iffues out of the fide of the hill, in the form of a Spring.

Dr. HALLEY was the inventor of this Theory. His attention was directed to the subject by the following fact. While busied in making some celestial observations, on a hill in the illand of St. Helena, he found, even when the fky was perfectly clear, that the quantity of vapour collected on his lenfes, every few minutes, was fo great, as wholly to impede his vision.

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Water, in the form of Vapour, is constantly rising from the Sea, in very large quantities, and in a state of freshness.

The following well-known facts may be adduced in

Rains are far more frequent and copious on mountains, than on plains; and in mountainous countries, than those that are level.

The earth on mountains is always moift; even during a drought.

Almost all springs iffue out of the sides of hills or mountains, or from lands adjoining them.

The loftiest mountains yield the most numerous riv-

ers; and the largest, also, where they are far enough from the fea.

Brooks are uncommon in champagne countries; and in countries, which have a stiff clay on the surface.

Brooks and Rivers may univerfally be traced to hills or mountains, when the month human necest and active the man first things which execute the R. W. Comen.

Streams, wells, and fountains are fullest in the Spring. In the Autum, many of them are absolutely dried up.

These facts collectively prove, that the quantity of water deposited on mountains is very great; that they are well qualified to retain it for the supply of Springs; and that springs and rivers are all of them in part, and many of them wholly, supplied from this source.

Various objections are urged against this theory. The first is that water will not soak far enough into the ground. M. De La Hire, a French Philosopher, to prove this, tried the following experiment. He dug a hole in the lower terrace of the Observatory at Paris; and placed therein, eight feet under ground, a large leaden bason, inclined a little towards one of its angles. To this was soldered a pipe, 12 feet long; which, after a considerable descent, reached into an adjoining cellar. He then silled up the hole with a mixture of sand and loam. After having kept the bason, in this situation, 15 years, (the ground being constantly exposed to all the rains and snows that fell,) he could never observe that a single drop of water had passed through the pipe into the bason.

On this subject, I will make two remarks. 1st. Water will certainly filter down, as easily as it will filter up; The attraction between the sand and the water is in each case the same; and in filtering down, it has the very serious affistance of Gravity. If then, it will filter upwards many miles, it will certainly filter downwards eight seet. 2dly. Our own eyes teach us, that deep wells frequently sail in dry weather; and that wells which are 30 or 40 feet deep are often very obviously raised by the hard rain of a single night. Hence I am led to conclude, either that the pipe employed in the experiment became clogged; or that Providence knows how to arrange earth, for the passage of water through it, bet-

ter than M. De La Hire.

It is likewise objected that springs often rise on plains, and sometimes on the summits of hills. Such occurrences are uncommon. Of the sew Springs which I

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have heard of as rifing on the tops of hills, almost all have been found, on examination, to issue many feet below the real apex. Were the fact otherwise, it would not be inconsistent with the Theory.—In stiff clayed grounds water may work its way a very considerable distance, before it finds an outlet. Springs, in wet lands, should be expected often to issue several miles from the refervoir. "And," in the language of Hutton, "if there happen to be a valley, between a mountain on whose top is a spring and the mountain which is to furnish it with water; the Spring must be considered as water conducted from a reservoir of a certain height, through a subterranean channel, to make a jet of an almost equal height."

It is also said, that some springs are not at all affected by drought. This may be owing to the capacity of the reservoir; and to the number of ducts which supply it,

and to the finallness of the drain.

But the principal objection is the infufficiency of va-

pour to supply the demands of springs and rivers.

Dr. Halley tried the following experiment, to determine the actual evaporation from the Mediterranean, so far as it is occasioned by heat. He filled a bason with brine, as falt as that of the Ocean; and heated it, over a pan of coals, to the temperature of the air in fummer. By a careful examination he found, that the quantity loft by evaporation was a tenth of an inch in 12 hours. He supposes the Mediterranean to be 40° long, and 4° broad; making a furface of 160 fquare degrees. According to the experiment, therefore, it will lose 5,280,000,000 tons of water in a day. The Mediterranean receives the waters of the following confiderable rivers; the Ebro, the Rhone, the Tyber, the Po, the Danube, the Neister, the Neiper, the Don, and the Nile. Dr. Halley supposed, that, on an average, each of these yields ten times as much water as the Thames; whereby he allowed for smaller rivers which fall into the fame fea. From an ingenious mensuration, he concluded, that the Thames discharges daily 20,300,000 tons.

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The nine rivers, therefore, discharge into the Mediterranean 1,827,000,000 tons of water in a day; which is but little more than a third of what is raised from its

surface in the fame time, by evaporation.

The above estimate of the water of the Thames is professedly overrated, and has since been found, by Mr. Dalton the Philosopher of Manchester, to be about one third two large. Reducing the nine rivers in this proportion, we shall find that their daily discharge is only 4,218,000,000 tons. This is rather less than 1-4th of the quantity evaporated from that sea; and leaves 4,062,000,000 tons, or 3-4ths of the whole, to meet the evaporation from that immense region, by which these

rivers and their branches are fupplied.

From a feries of observations respecting the annual supply of rain, made at 31 different stations in England, during different periods of from 1 to 21 years, Mr. Dalton sinds the average quantity for the whole of England and Wales to be 31 inches. To this he adds 5 inches for dew; making a total of 36 inches or three feet. After a minute examination of the rivers of those two countries, he estimates their whole annual discharge to be nine times that of the Thames. Allowing this calculation to be correct, the rivers will exhaust only 13 inches of rain; leaving 23 inches to evaporate from the land: a residuum, which he finds, by a number of experiments, to be amply sufficient.

The quantity of rain, which falls in our own country, is believed to be confiderably greater than that in any country of Europe. The exact amount however cannot, at prefent, be flated with precision; as few gentlemen have hitherto made the requisite observations. I can only observe, that I have seen various registers of rain kept at different places for considerable periods: and the average results of a still larger number. From these I am led to conclude, that the quantity of water, which annually falls in rain snow and hail, will average 45 inches. To this should be added 5 inches, at least, for dew; making the whole 50 inches. The United

States comprise a million of square miles. Thirty five cubic feet of rain water weigh a ton. A mass of water, covering 1,000,000 square miles to the depth of 4 feet 2 inches, contains 116,160,000,000,000 cubic feet;

amounting to 3,318,860,000,000 tons.

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By an examination of the Map of the United States it will appear, that, on its borders and within it, there are 26 great primary rivers, or rivers communicating directly with the Ocean These are the St. Lawrence. Penobscot, Kennebec, Ameriscoggin, Saco, Merrimac, Connecticut, Hudson, Delaware, and Susquehannah, in the North; and the Potowmac, Rappahanoc, York, James, Roanoke, Pamlico, Neufe, Cape Fear, Pedee, Santee, Savannah, Ogeechee, Altamaha, Apalachicola, Mobille, and Missifippi in the South. The St. Lawrence receives half of its water from the Canadas; and more than three fourths of the fupplies of the Miffiffippi are furnished by Louisiana. Of this the proof is direct. The Obio, the only very large tributary from the east, is smaller than the Arkansas; and the Illinois, the only remaining eaftern branch worthy of notice, is much less than the St. Francis. On the West, however there are also the Missouri, which is larger than the Misfiffippi itself before the junction, added to all the branches from the east; and Red River, which is much larger than the Ohio.

Of the remaining primary rivers, the Amerifcoggin, Saco, Merrimac, Rappahannoc, York, Pamlico, Neufe, and Cape Fear, are about as large as the Thames; and the Penobscot, Kennebec, Ogeechee, and Altamaha, are not much larger. Taking these facts into consideration, I am fatissied, that, if we admit each of the 26 primary rivers to yield ten times as much water as the Thames, we shall allow a sufficient overplus to supply all the primary streamlets. On this supposition the annual discharge of water from all the rivers of the Union will amount to 1,281,660,000,000 tons to supply the demands of

evaporation from the land.

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I am aware, that fuch calculations, in their own nature, do not admit of that certainty which is demanded in demonstration. Still they are sufficiently accurate to leave the mind satisfied. I am also aware, that their want of certainty comes with an ill grace from the mouth of the objector. As the advocates of the theory prove, in the outset, that many springs are wholly supplied in this manner; the case with respect to the others, is prima facie with them. As the objector takes the issue, the burden of proof rests upon his shoulders. Until it is actually surnished I believe we shall all admit, that Springs owe their origin only to Vapour.

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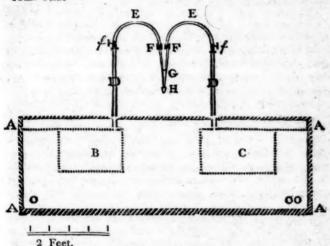
No. XX.

EXPERIMENTS OF THE PARTY OF THE

On the Fusion of various refractory Bodies, by the

By BENJAMIN SILLIMAN, PROF. CHEM. AND MIN. IN YALE-COLLEGE.

A SECTION OF THE PNEUMATIC CISTERN OF YALE COLLEGE WITH THE COMPOUND BLOW PIPE OF MR. HARE FOR BURNING HIDROGEN, MINGLED WITH OXIGEN GAS.



REFERENCES TO THE FIGURE.

AAAA.—The pneumatic cistern, filled with water; for a plate, and full description, see the Boston Edition of Henry's

Chemistry.

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B.—A Gas Reservoir, of the capacity of twelve gallons, filled with oxigen gas, either by the action of the hydrostatic bellows at O, or, by a recurved tube, passing from above, through the water, and hooked under B: parallel, and contiguous to B, on the other side of the cistern, is another gas reservoir, of the same capacity, which may be connected with B, or not, at pleasure.

C.—The same, in every respect, only C is filled with hidrogen, by hydrostatic bellows at OO, or by a recurved tube, as above.

D.—Copper Tubes, half an inch in diameter, furnished with stop cocks at f, and inserted into the gas reservoirs B. C.

E.—Recurved tubes of flexible metal, furnished with double screws at F, which connect them with a pair of brass blow pipes, cut off at G, and soldered to two strong cast silver tubes, which screw, air tight, into H, an inverted pyramidal piece of platinum, in which, two converging ducts as large as a pin are perforated, forming a continuation of the tubes, and uniting in a common passage, somewhat larger, just before their exit, at the common orifice below. The subject to be operated upon is sustained by charcoal, or forceps, and held by the hand, just below the orifice in the piece H.

e:

The gasses at B. C. are under hydrostatic pressure, which is easily recruited, as the gasses run out, either by throwing common air with the bellows, into one of the spare reservoirs, or, by introducing more of either of the gasses into the appropriate reservoir, and, peculiarly of hidrogen, both, on account of the facility with which it is obtained, and because, twice as much of it, in bulk, is wanted, as of oxigen.

The rapidity of efflux of the gasses, and their due proportion, is easily regulated, by turning, more or less, the keys of the stop cocks at f, and the effects of either gas alone, may be observed, by shutting the stop cock leading to the other.

When the compound flame is desired, the hidrogen is first let out, and fired; the blaze should be somewhat larger than that of a candle; the oxigen is then let into the hidrogen till the effect is the greatest, which a little habit will soon ascertain.

The flame of the hidrogen is very much narrowed, by the introduction of oxigen, and there is no appearance of peculiar splendor or heat, till some body, capable of reflecting the light and heat, is placed in the focus, which is usually about

one fourth of an inch below the orifice.

All the apparatus below **F F** is easily detached, by turning the double screws;—the strong silver tubes are intended to prevent fusion of this part of the apparatus, and to admit of connexion with the platinum piece by means of a screw cut on the silver tubes; this obviates the necessity of using a solder, which would be very liable to melt, and, the platinum piece is, for a similar reason, substituted for the silver cylinder originally used by Mr. Hare, as experience has shewn that these are liable to fusion.

No flux or addition of any kind was employed in the following experiments.

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ON THE FUSION OF VARIOUS REFRACTORY BODIES BY

THE COMPOUND BLOW-PIPE OF MR. HARE.

THE philosophical world behold with pleasure and astonishment, the effects produced on the fusion and combustion of bodies, by a stream of oxygen gas, directed upon burning charcoal. The fplendor of these experiments arrefted universal attention, and Lavoisier, with his gazometer, was enabled, in this manner, to produce a degree of heat, furpaffing that of the most powerful furnaces, and even of the folar focus. Bodies which no degree of heat, previously applied, had been able to foften, now became fluid, and philosophy appeared to have attained the limit of its power in exciting heat; indeed, it feemed to have advanced, very far, towards realizing the opinion, that folidity and fluidity are accidental attributes of bodies, dependant folely on the quantity of caloric which they contain, and that therefore, they may be supposed capable of existing in either of these conditions.

Still however, there were, in fact, many important exceptions. Of the primitive earths, Lavoisier had been enabled to fuse only alumine—while the rest remained refractory, and seemed fully entitled to the character of infusibility, usually attributed to this class of bodies: many native minerals and especially those which are most distinguished for hardness, beauty, and simplicity of composition, maintained the same character, and some of them resused to melt even when heat-

ed with powerful fluxes.

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The beautiful invention of Mr Robert Hare of Philadelphia, by which he fucceeded in burning, with fafety and convenience, the united stream of oxygen and hidrogen gases, greatly extended our dominion over refractory bodies, and presented new and very interesting results. Mr. Hare's memoir, originally communicated to the Chemical Society of Philadelphia, has been

fome years, before the public, and has been republished and handsomely noticed, both in France and England. Still however, his results have not found their way into the Systematical books on Chemistry, (with the exception of Mr. Murrays system,) notwithstanding that some of the European Professors have availed themselves of Mr. Hare's invention, so far as to exhibit his most splendid and striking experiments to their classes.

The writer of this article, although fully disclaiming any share in Mr. Hare's invention, was early associated with him in his experiments; they excited in his mind a degree of interest, which led him to hope that they would be repeated and extended by others, but, as nothing of this kind has appeared in this country, perhaps the following experiments may not be altogether uninteresting, especially as they were performed with an apparatus, of a construction somewhat more simple than the original.

It will be necessary to recollect that Mr. Hare not only melted alumine, which Lavoisier had done before, but also silex and barytes, and, by subsequent experiments, he added strontites, to the list of suspenses: he was inclined to believe that he had volatilized gold and silver, a conclusion which was rendered highly propable by his having afterwards evidently volatilized platinum.

The experiments of Mr. Hare, as will appear below, have been repeated by the writer of this paper with fuccess, and many other bodies among the most refractory in nature, have been melted. For the sake of shewing how far the experiments now to be recited have affected our knowledge of the dominion of heat, quotations, for comparison, will occasionally be made, from one of the latest and most respectable chemical authorities.

(Murray's System 2d Ed.)

BODIES SUBMITTED TO THE HEAT OF THE COMPOUND BLOW-PIPE OF MR. HARE.

PRIMITIVE EARTHS.

SILEX—being in a fine powder, it was blown away by the current of gas, but when moistened with water, it became agglutinated by the heat, and was then perfectly fused into a colourless glass.

ALUMINE—perfectly fused, into a milk white enamel.

BARYTES—fused immediately, with intumescence, owing to water, as observed by Lavoisier; it then became solid and dry, but soon melted again into a perfect globule, a greyish white enamel.

STRONTITES—the same.

GLUCINE- perfectly fused into a white enamel.

ZIRCON—the fame.

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LIME-in small pieces, it was immediately blown off from the charcoal; to prevent this, as well as to obviate the suspicion, that any foreign matter had contributed to its fusion, the following expedient was reforted A piece of lime, from the Carrara marble, was strongly ignited, in a covered platinum crucible; one angle of it was then shaped into a small cylinder, about one fourth of an inch high, and fomewhat thicker than a great pin: the cylinder remained in connection with the piece of lime; this was held by a pair of forceps, and thus the small cylinder of lime was brought into contact with the heat, without danger of being blown away, and without a poffibility of contamination; there was this farther advantage, (as the experiment was delicate and the determination of the refult might be difficult,) that, as the cylinder was held in a perpendicular polition, if the lime did really melt, the column must fink and become, at least to a degree, blended with the supporting mass of lime. When the compound flame fell upon the lime, the fplendor of the light was perfectly insupportable, by the naked eye, and when viewed through deep coloured glasses (as indeed all these experiments ought to be,) the lime was seen to become rounded at the angles, and gradually to sink, 'till, in the course of a few seconds, only a small globular protuberance remained, and the mass of supporting lime was also superficially fused at the base of the column, through a space of half an inch in diameter. The protuberance, as well as the contiguous portion of lime, was converted into a perfectly white and glistening enamel; a magnifying glass discovered a few minute pores, but not the slightest earthy appearance. This experiment was repeated several times, and with uniform success; may not lime there-

fore be added to the lift of fufible bodies?

MAGNESIA.—The fame circumstances that rendered the operating upon lime difficult, existed, in a still greater degree, with respect to magnesia; its lightness and pulverulent form rendered it impossible to confine it for a moment upon the charcoal, and as it has very little cohesion, it could not be shaped by the knife as the lime had been. After being calcined, at full ignition, in a covered platinum crucible, it was kneaded with water, 'till it became of the consistence of dough. It was then shaped into a rude cone as acute as might be, but still very blunt; the cone was three fourths of an inch long, and was supported upon a coiled wire.

The magnesia, thus prepared, was exposed to the compound flame: the escape of the water caused the vertex of the cone to fly off in repeated flakes, and the top of the frustrum, that thus remained gave nearly as powerful a reflection of light as the lime had done; from the bulk of the piece (it being now one fourth of an inch in diameter at the part where the flame was applied) no perceptible finking could be expected. After a few feconds, the piece being examined, with a magnifying glass, no roughnesses or earthy particles could be perceived on the spot, but a number of glaffy, fmooth protuberances, whose furface was a perfectly white enamel; this experiment was repeated with the fame fuccefs. May not magnelia, then, be also added to the table of fusible bodies?

YTTRIA—was the only remaining primitive earth, but no specimen of it could be obtained.

Perhaps then we shall be justified in saying, in future, that the primitive earths are suffile bodies, although not suffile in surfaces, in the solar socus, nor, (withthe exception of alumine, and possibly, barytes,) even by a stream of oxygen gas directed upon burning charcoal.

PLATINUM—was not only melted but volatized with firong ebullition.

VARIOUS MINERALS.

ROCK CRYSTAL,—transparent and colourless. This mineral was instantly melted into a beautiful white glass. "It not only does not melt in the focus of the most powerful burning mirror, but, it remains without fusion, at least when in the state of Rock Crystal, in the still more intense heat, excited by a stream of oxigen gas directed on burning charcoal." (Murray II. 261.) "It is even imperfectly softened by the intense heat, excited by a stream of oxigen gas, directed on the slame." (of the blow pipe lamp.)—(Ibid III. 513.)

COMMON QUARTZ—fuled immediately into a vitreous

globule.

GUN FLINT-melted with equal rapidity; it first became white, and the fusion was attended with ebullition and a separation of numerous small ignited globules which seemed to burn away as they rolled out of the current of slame; the product of this susion was a beautiful splendid enamel.—" It is insusible before the blow pipe but loses its colour."—(Ibid. 518.)

CHALCEDONY—melted rapidly, and gave a beautiful bluish white enamel resembling opal. "It is infusi-

blebefore the blow pipe."—(Ibid. 516.)

ORIENTAL CARNELIAN—fufed with ebullition, and produced a femitransparent white globule with a fine luftre.

RED JASPER—from the Grampians, was slowly fused with a sluggish effervesence, it gave a greyish black slag, with white spots.

"It is infusible before the blow pipe, even when the flame is excited by a stream of oxygen gas." (Ibid, 519.)

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SMOKY QUARTZ-or fmoky topaz melted into a col-

ourless globule.

Beryl—melted instantly, into a perfect globule, and continued in a violent ebullition, as long as the slame was applied, and when, after the globule became cold, it was heated again, the ebullition was equally renewed; the globule was a glass of a beautiful blue, is milky white.

"The beryl is melted with difficulty before the blow-

pipe alone, but eafily when borax is added."

EMERALD of Peru. (Ibid 511.)

The fame, only the globule was green, and perfectly

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OLIVIN—fused into a dark brown globule, almost black, "It can scarcely be melted by the blow pipe without addition." (Ibid 534.)

VESUVIAN—instantly melted into a beautiful green glass.

"It melts before the blow pipe into a yellowish glass."

(Ibid 534.)

LEUCITE—instantly sused into a perfectly transparent white glass; the sussing ebullition, and many ignited globules darted from it and burnt in the air, or rolled out upon the charcoal and then burned. Were they not potassium? This stone contains sull 20 per cent of potass; this hint will be resumed below.

" It is not fused before the blow pipe." (Murray III.

534.)

CHRYSOBERYL—(Cymophane of Hauy) was immediately fused into a greyish white globule. "It is not melted by the blow pipe," (lbid 499.)

A CHRYSTALIZED MINERAL.

From Haddam, Connecticut, according to the Abbe Hauy it is Chrysoberyl, according to Col. Gibbs, Corundum: it fused with ebullition, and scintillations, and produced a very dark globule almost black.

Topaz—of Saxony, melted with strong ebullition, and became a white enamel. "It is infusible before the blow pipe, but melts when borax is added." (Ibid 498.) Sappar or Kyanite—perfectly and instantly sused, with ebullition, into a white enamel.

"It remains perfectly unaltered before the flame of the blow pipe even when excited by oxygen gas."

(Ibid 499.)

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CORUNDUM-of the East Indies, was immediately and

perfectly fused, into a grey globule.

CORUNDUM—of China, the same with active ebullition.

Corundum "is not fused by the slame of the blowpipe on charcoal even when soda or borax is added
to it." (Ibid 495.)

ZIRCON—of Ceylon melted, with ebullition, into a white enamel. "It is not melted alone before the flame of the blow pipe, but if borax is added it forms a trans-

parent glass." (Murray III. 539.)

HYACINTH-of Expailly fused into a white enamel.

"It loses its colour before the flame of the blowpipe, but it is not fused; it melts with borax into a transparent glass." (Ibid 540.)

CINNAMON STONE—instantly fused into a black globule

with violent ebullition.

Spinelle Ruby—fused immediately into an elliptical red globule. "It does not melt before the blow pipe but is fused by the aid of borax. [Ibid 497.]

STEATITE—melted with strong ebullition into agreyish flag.—" It does not melt before the blow pipe, but

becomes white and very hard." [Ibid 482.]

Porcelain, common pottery, fragments of Hessian crucibles, Wedgwood's ware, various natural clays, as pipe and porcelain clay, fire and common brick, and com-

pound rocks, &c. were fused with equal ease.

During the action of the compound flame upon the alkaline earths, provided they were supported by charcoal, distinct globules often rolled and darted out from the ignited mass, and burnt, sometimes vividly, and with peculiarly coloured slame. From the nature of the

experiments, it will not be eafy to prove, that these globules were the basis of the earths, and yet there is the strongest reason to believe it; circumstances could scarcely be devised, more favourable to the simultaneous sufficient and decomposition of these bodies; charcoal highly ignited for a support and an atmosphere of hydrogen also in vivid and intense ignition; that the oxygen should be, under these circumstances, detached, is not surprising, but the high degree of heat, and the presence of oxigen necessarily burn up the metalloids almost as soon as produced. If means could be devised to obviate this difficulty, the blow pipe of Mr. Hare might become an important instrument of analitical research.

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We can fearcely fail to attribute fome of the appearances, during the fusion of the Leucite, to the decom-

position of the potash it contains.

This impression was much strengthened by exposing potash and Soda to the compound stame, with a support of charcoal; they were evidently decomposed: numerous distinct globules rolled out from them, and burnt with the peculiar wivid, white light, and stash, which these melalloids exhibit, when produced and ignited in the galvanic circuit. It is hoped that these hints may induce a farther investigation of this subject.

The experiments which have now been related in connection with the original ones of Mr. Hare, sufficiently shew that science is not a little indebted to that gentleman for his ingenious and beautiful invention.— It was certainly a happy thought, and the result of very philosophical views of combustion, to suppose that a highly combustible gaseous body, by intimate mixture with oxygen gas, must, when kindled, produce intense heat: and it is, no doubt, to this capability of perfectly intimate mixture, between these two bodies, that the effects of the compound blow pipe, are, in a great measure, to be ascribed.

This communication has already been extended farther than was contemplated, but on concluding it, it may be allowable to remark, that there is now, in all

probability no body, except fome of the combustible ones, which is exempt from the law of fusion by heat. If the primitive earths, and fuch minerals, as several of those which have been mentioned, above, are fusible, no doubt can be entertained that all other mixtures and combinations of earths are fufible also: for, fuch mixtures and combinations are known to be more fufible than the primitive earths; the metals are more fufible than the earths, and the diamond along with carbon in its other pureft forms, appears to be really the only exception; and it is probable that this is only a feeming one, for, it is scarcely possible to expose these bodies to the heat of the compound blow pipe, without at the fame time burning them up: could the heat be applied without exposing them to the contact of oxigen is it not probable that they also would be added to the lift of fufible bodies? many could previously to have

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Yale College, May 7, 1812.

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No. XXI.

OBSERVATIONS

On the Comet of 1811.

By JEREMIAH DAY, PROF. OF MATH. AND NAT. PHILO. IN YALE COLLEGE.

ON the 9th of September last, I began a course of observations, on the Comet which has lately withdrawn from our view. They were continued, with considerable intervals of interruption, till the 13th of January. Though I was possessed of no instrument, which would give the place of the Comet, with the utmost precision; yet the observations were sufficiently correct, for a determination of the elements of its orbit, with a near

approach to exactness.

Few Comets have prefented themselves to our view. under circumstances more favourable, for observing their motions. Many have appeared for a few days, and then, fuddenly retired. But this was visible, for feverable months, even to the naked eye. During this period, it traversed a space, of more than 130 degrees, in the heavens. A general idea of the direction of its motion may be obtained, by conceiving a great circle to be drawn through the flar , in the extremity of the tail of the Great Bear; and the bright ftar in the Eagle. It was near the first of these, on the 2d of October; and very near the latter, on the 1st of December, as Mr. Bowditch had predicted, two months before. It passed across the constellation of the Great Bear, the head of Asterion, the right arm of Bootes, the northern part of Hercules, the Eagle, and the bow of Antinous; and when I faw it last, on the 13th of January, it was a little fouth of the head of Equuleus. Clouds intercepted the view of it, one or two of the following evenings; then fucceeded a period of moonlight; and, when the moon had passed the full, the

comet had advanced too near the fun, to be longer visible. I had found its place, from time to time, by taking its distance from some of the principal fixed flars, with a fextant, graduated to every ten feconds. The time was given, by a well regulated clock. To diminish the errours of observation, each of the distances was generally measured ten or twelve times in fuccession, and a mean taken from the whole. The requisite corrections, for the refraction of the atmosphere, were afterwards applied. The following are fome of the observed distances. Two or three of the first are not probably very correct, as they were taken, when the comet was fo near the horizon, as to render the

riem of it obscure and the refraction uncertain

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	17	8	52	2	7	20	0	32	22	39	
	18	9	3	2	5	47	12	33	24	24	
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Nov.	3	7	46	. 1	1	53	4	30	19	4	
	49d	7	50	1	2	17	18	28	53	12	
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This comet came to us, from the regions of the fouth. It croffed the ecliptic on the 11th of July, and advanced fo far to the north, that on the 28th of September, it was within the circle of perpetual apparition of this latitude, and, of courfe, continued above the horizon, the whole 24 hours. In two or three days from this, it reached its greatest northern declination, about 50 degrees; and then began to return towards the equator. But it did not attain its highest geocentric latitude, till the middle of October. Its apparent place was then, within 28 degrees of the pole of the ecliptic. Itsmotion, as feen from the earth, was, during the month of September and October, nearly in the arc of a great circle. But the latter part of the time, in which it was visible, it sensibly deviated to the east of its former direction; though, on account of its moderate angular velocity, its elongation from the fun was then rapidly diminishing. The rate of its apparent motion has been less, than that of many other comets. feen in 1472 moved 40 degrees, and another in 1770, 45° in 24 hours. But the progress of the late comet never exceeded two degrees in 24 hours: and during the latter part of the time, it did not amount to half a degree.

This is a general view, of the apparent motion of the comet, to a spectator on the earth. But its real motion will be found to be widely different: so much so, that, while the apparent motion was, most of the time, towards the east; its real motion in longitude was westward, or contrary to the order of the signs. To determine the direction and rate of this motion, it was necessary to refer it to the sun, as a centre, and to calculate the elements of the comet's orbit. These are the Perihelion distance, the time of passing the perihelion, the longitude of the perihelion, the longitude of the node, and the inclination of the orbit to the ecliptic. To obtain a first approximation, to the time and distance of the perihelion, some of the early observations were made use of. But for the sinal corrections, it was ne-

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ceffary to take fuch, as had a greater interval of time between them. The observations of the 17th of September, the 17th of October, and the 14th of December were accordingly felected: and the elements were corrected by the method of La Place, as given, in Sir Henry Englefield's treatife on the "Determination of the orbits of Comets." To fucceed with this method, it was necessary that the elements should be very carefully prepared. For the convenience of calculation, the three observations were reduced to 8 o'clock, mean time at Yale College, by applying to the observed diftances, a proportional part of the diurnal variation. The right ascension and declination of the stars, from which the distances of the comet had been taken, were found, by comparing the catalogues of Dr. Bradley, De La Caille, Piazzi, Zach, Wollaston and Pond, in Vince's Aftronomy, Hutton's Dictionary, Rees' Cyclopedia, and Mendoza's Tables. These generally agreed, within a very few feconds, with respect to the ftars in question. The corrections, for precision, aberration, and nutation were applied; and from the right ascensions and declinations, the latitude and longitude of the stars, were calculated to feconds; as were also the geocentric latitudes and longitudes of the comet. The latter were farther corrected, by applying the equations, for the aberrations of light. The three latitudes and longitudes thus obtained, were,

 Geocen. Lon.
 Latitude.

 September 17,
 153° 64′ 36″
 38° 39′ 56″

 October 17,
 221 54 28
 62 27 43

 November 14,
 286 9 15
 41 50 53

Taking these, and the approximate perihelion distance and time, as the basis of the succeeding calculation, I proceeded to derive from the comet's observed longitudes and latitudes, the arc which it described, between the first and second observation; and also between the second and third; for the purpose of comparing them, with the corresponding arcs, obtained by taking the difference of the three anomalies. If these had been

found to agree; no other correction of the elements would have been requifite, to adapt them to the three observations. But this was not to be expected, on the first trial. There was a difference of several degrees. It was necessary therefore to make a second, and a third hypothesis, in one of which, the perihelion distance was changed, and, in the other, the time of perihelion. From the errours of the three, the equations were formed, for determining the corrections to be applied to the assumed elements. After two sets of hypotheses, a perihelion distance, and time of perihelion were obtained, which gave the arc described by the comet, between the 17th of September and the 14th of November, and measured by an angle at the fun, within less than half a minute of that deduced from the observed longitudes and latitudes:

The one being The other 58° 26′ 14″ 58 25 52

Difference

00 0 22"

As this difference is within the limits of the unavoidable errours of observation, I did not attempt to carry the approximation any farther. Some slight variations would probably have been made, by processes depending on a comparison of all the observations. But as they would be of little amount; and as perfect accuracy would not be attainable after all; I proceeded to calculate the remaining elements, from the two already found. The whole together were as follows:

Perihelion distance 1.0329; the mean distance of the

fun from the earth being 1.

Time of perihelion Sept. 12th, 1 o'clock, P. M.

mean time at Greenwich.

Longitude of the afcending node,
Inclination of the orbit to the ecliptic,
Longitude of the perihelion, counted
on the orbit,

s.

4 20° 22'
73° 4'
2 15° 14'

Distance of the perihelion from the node, 2 5° 8'

Motion retrograde.

The following distances of the comet from the earth, and the fun were also calculated:

stances from the	earth.
earth's orbit.	In miles.
2.411	229,045,000
1.512	143,670,000
1.224	116,290,000
1.591	151,145,000
2.359	224,070,000
	2.411 1.512 1,224 1.591

Distances from the Sun. Semidiam. of earth's orbit. Sept. 12th 1.0329 98,125,500 Sept. 7th and 17th, 1.0367 98,490,000 Sept. 2d and 22d, 1.0470 99,470,000 Aug. 23 and Oct 1, 1.081 103,000,000 Aug. 8 and Oct. 17, 1.189 113,000,000 July 23 and Nov. 1, 1.323 126,000,000 July 11 and Nov. 14. 1.463 139,000,000

June 23 and Dec. 1,

157,000,000 June 9 and Dec. 16, 1.831 174,000,000 May 23, 1811, and 7 2.022 192,000,000 Jan'y 1, 1812,

1.655

April 22, 1811, and 2.385 227,000,000 Feb. 1, 1812,

In addition to the elements given above, one thing farther is necessary to complete the theory of the comet's motion—the period of its revolution. This is an article on which much labour has heretofore been bestowed, with very little fuccefs. Two methods are obvioully fuggested, for determining the time of a comet's The most direct of these, is to derive, by calculation, the figure and dimensions of the whole orbit, from that fmall part of it, which is described, while the comet It refults from the well known laws of gravitation, that any body, moving round the fun, and influenced by the attraction of no other body, must move in one of the three conic fections, the ellipsis, the parabola,

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If it is ever to return, in a regular or the hyperbola. orbit, it must revolve in an ellipsis. If its motion is in either of the two other figures, after having once paffed the perihelion, it will continually recede from the fun, and will return no more; unless its direction should be changed, by the attraction of some other body. In each of these cases however, if a portion of the path of a comet could be obtained by observation, with perfect exactness; from this might be deduced, the remaining parts of the orbit, on the supposition, that its figure should remain unaltered. But here two insuperable difficulties occur, one from the errours of observation, the other from the liability of the comet, to be diverted from its course, by the attraction of the planets, and perhaps of the fixed stars. The ground of these difficulties, lies not in the processes for calculating the orbit; but in taking the observations, and in the change of the orbit itfelf. Errours too minute to be avoided, even by the nicest instruments hitherto in use, might be sufficient, in certain cases, especially where the orbit is very eccentric, to make a variation of many years, in the periodical time. The most able computers, have accordingly differed whole centuries, in the periods, which they have respectively assigned to the same comet. The only cafes to which calculation can be applied, with any hope of fuccess, are those in which the time of revolution is very fhort. The comet of 1770, has been supposed to be one of this class. Lexell, Pingre, and Burckhart all agree in giving it a period of about five years and an half. There is reason to believe, that this is the orbit, which really corresponds with that part of its motion, which was observed. But notwithstanding this, the comet has never been feen fince; though it ought to have returned fix or eight times, in the intervening forty years.

If its orbit was truly afligned, it feems it must have been since altered, by the attraction of the planets or of some other body, whose influence may be sufficient, not only to vary materially the time of revolution; but

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even to change the comet's path, from an ellipsis, to a parabola or hyperbola, so that it shall never return.

The fecond method, which has been adopted, for afcertaining the periodical time of a comet; and, on which confiderable reliance has been placed, is, to compare the elements which are commonly computed, with those of all the other comets, on which calculations have been If feveral of them are found to have the fame elements; it is concluded they must be one and the fame comet. For it is scarcely credible that in the endless variety which is found to prevail, in the directions and rate of their motions, there should be any two, which should have precisely the same perihelion distance, the fame longitude of their nodes, the fame place of perihelion, and the same angle of inclination to the ecliptic: and, in addition to all this, that they should appear at intervals of time fo regular, as to correspond with the fuccesfive returns of the fame comet. On this ground, it has been supposed, that the comet of 1759 has a period of about 75 years, because one with similar elements was feen in the years 1456, 1531, 1607, and 1682. But no fuch correspondence is found, in the case of the late comet. From a comparison of its elements with others whose motions have been heretofore observed and calculated; it will be feen that this is one, which is not included in the lift. In perihelion diffance, it nearly agrees with one, which appeared in 1718; but differs from it, more than 40 degrees, in the inclination of the orbit. In the longitude of the node, it is within lefs than a degree of that in 1759; but varies materially from it, in the longitude and distance of the perihelion. In the inclination of the orbit, it differs but little from one in 1097 and another in 1763; but has no agreement with them, in the other elements. So that this method of determining the periodical time fails, in its application to the present case. And even, if the elements of this comet, had been found to agree, with those of any preceding one; it would have ferved rather to flew, in what interval of time it has returned, than to give us affurance, that it will return, at the same interval, hereafter. The reasons of this, have already been stated to the Academy, by Col. Mansfield, in his ingenious re-

marks, on the comet of 1807.

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The motion of the late comet, as feen from the fun. was nearly north and fouth; inclining however, about 17 degrees, towards the northwest and southeast. It passed northward, through the constellations of the Lion, the Great Bear, and the Camelopard. On the 19th of September, to a spectator in the sun, it was within three degrees of that part of the heavens, to which the axis of the earth is directed. It reached its greatest northern heliocentric latitude 73 degrees, on th 1st of October; and then returned, through Cepheus, the left wing of the Swan, Equuleus and Aquarius. It will proceed through the Microscope, the Indian, and the Octant; till, at its greatest distance from the earth, it will be between the Chameleon and the Flying-fifh, within 17 degrees of the fouth pole of the ecliptic. Should it vifit us again, it will return through the Ship, the Air-pump, Hydra and the Sextant.

This comet has not approached so near to the fun, as most of those, whose elements have been calculated. Its least distance, is 98 millions of miles; a little greater than that of the earth from the fun. Its path lies between the orbits of the earth and Mars. It is the opinion of Newton, that no comet is ever feen by us, when farther distant than Jupiter. Of about one hundred, whose elements have been calculated, all, except four, have come nearer to the fun than Mars. These four fell between Mars and Jupiter. About 20 came between Mars and the earth; 15 between the earth and Venus; nearly 30 between Venus and Mercury; and about the fame number, within the orbit of Mercury. The remarkable one of 1680, came much nearer to the fun, than any other: fo near, as to be heated, according to Newton, 2000 times hotter than red hot iron. Its least distance, from the surface of the sun, was not equal to

a fifth part of his diameter.

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There is no danger, that the late comet will ever firike the earth; unless its orbit should, hereafter, be materially changed. Its least distance, in the present instance was more than an hundred million miles. This was about the middle of October. Its nearest approach to the *orbit* of the earth, was on the 11th of July, about 40 million miles. It can never come much nearer than this, without a change in the position or figure of its orbit. Its rate of motion, when nearest the sun, was 95,000 miles in an hour. This is a velocity, 120 times

greater than that of found, or a cannon ball.

For the purpose of determining, if practicable, the size of the comet, I viewed it feveral times, through a three feet reflecting telescope, with a magnifying power of 140. But I was unable to perceive any nucleus, with a disk sufficiently defined, to admit of measuring its diameter. This will not appear furprifing, when it is confidered, that out of 15 or 20 comets, which Dr. Herschel has had an opportunity of observing, there were only two or three, on which he was able to discover any regular disk, even with the very powerful telescopes in his possession. For want of a proper regard to the distinction between the different parts of a comet; no great dependence is to be placed, upon the accounts given us, of the fize of those, which were formerly feen. They have been frequently reprefented, as larger than any of the planets; and, in fome instances, as appearing nearly equal to the fun and These statements may be true, if they are meant to refer to the whole of the luminous spot, or body of light, which is perceived, by the naked eye. But this, in many cases at least, consists of three parts -the nucleus, the head, and the coma. The nucleus appears to be a compact fpherical body, with a circumference regularly defined, like a planet. The head, is a body of denfe light, which, like an atmosphere, furrounds the nucleus. The coma, is a space occupied by a fainter light, extending confiderably farther round, than the head. All these may be so blended, to the view of the naked eye, as to appear to conftitute but

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one body of light; fo that the diameter of the coma may be taken, for the diameter of the comet itself. A telescope, by expanding the head, often makes it appear less bright, and less distinct. It spreads it out, into a kind of nebula, which has no well defined boundary; but which may, notwithstanding, be mistaken for the nucleus. The latter is generally too small, to be distinctly perceived, by ordinary telescopes; and in many inftances, cannot be discovered, by those of the greatest magnifying power. That of 1807, according to Dr. Herschel, fubtended an angle of only one or two feconds; while the diameter of the coma was two or three hundred times as great La Place and others have even advanced the opinion, that the whole body of the comet is fometimes a mere collection of aeriform fluids, most condensed near the centre, but containing no fubftance fufficiently compact, to obstruct entirely the passage of light. In the present instance, I observed nothing, which could either invalidate or confirm, fuch an hypothesis. The head appeared, like an obscure cloud or mift. The coma was nearly of the fame brightness with the tail. A darker space, or a zone of weak light, very perceptibly fainter than either the head, or the coma, intervened between the two.

The tail of this comet, was more splendid, than is common: though some others have exceeded it, both in brightness and extent. That of 1680, was two or three degrees in breadth, and about 70, in length. That of 1759 was, according to Pingre, 90° long. One in 1618 is faid to have extended more than a hundred

The length cannot be measured, with exactness. For the extremity does not terminate at once; but gradually diminishes in brightness, till it is lost, in the still fainter light of the sky. The dimensions will vary, according to the state of the atmosphere. They will even appear different, to different eyes, at the same time. About the middle of October, I judged the tail of this comet to be nearly or quite 15 degrees long. It could not be less than twelve. If it be taken at 12° on the 17th;

and the breadth at half a degree; it will be found, by calculation, that the length in miles, was 40,000,000; and the breadth 1,000,000: the whole occupying a fpace, which would not be filled by 60 million fuch

globes as our earth.

Stars were frequently to be feen through the tail; though they were confiderably obscured. That side of it, which was turned towards the part of the heavens, to which the comet was moving, was a little convex, as usual, and brighter than the opposite side. I repeatedly observed a dark line, like a shadow, extending from one end of its tail; to the other; and passing nearly through the middle, but a little further from the advancing fide, than from the other. The whole appearance was fuch, as to correspond very well with the fupposition, that the tail was hollow. The light was denfest on each side, and gradually diminished, towards the middle, where was a narrow space, almost as dark as the neighbouring parts of the fky. This peculiarity has not been mentioned, in the accounts of other comets, except in a few inftances. Hevelius states, that one' which he faw in 1665, cast a shadow through the middle of the tail. A fimilar appearance was observed in the comet of 1744; and also by Cassini, in that of 1680.

I have not entered into any speculations on the nature and use, of this wonderful train of light, which is as unaccountable, to the astronomer, as to the vulgar obser-Some extravagance of conception is certainly excufable, in attempting to explain the constitution of a luminous object, which occupies a greater space, than all the other bodies in the folar fystem. But the schemes which have hitherto been proposed, for this purpose, are rather to be confidered as displays of the power of imagination; than specimens of the exercise of sound and fober reason. Those who have a taste for these visionary hypotheses, may easily contrive them for themselves; or may find, in the common aftronomical works, a very convenient affortment of them, adapted to the fancy, of almost every description of readers.

Yale-College, March 20th, 1812.

CALCULATION

Of the Longitude of Yale-College, from the Solar Eclipse of September 17th, 1811.

By JEREMIAH DAY AND JAMES L. KINGSLEY,

PROFESSORS IN YALE-COLLEGE.

IT is important, that the Latitude and Longitude of places, especially of those in which astronomical observations are to be made, should be settled, with as much

accuracy, as the nature of the case will admit,

The Latitude is eafily obtained, from the meridian altitudes of the fun, and the fixed stars. But some more complicated process is necessary, for determining the Longitude. Most of the methods in use, for this purpose, depend on the motion of the moon. in her orbit. Several of them, however, cannot be relied on, to give the longitude, with any great degree of accuracy. That founded on the observation of folar eclipses, has an advantage in this respect, over most of the others. As we had feen no circumitantial account of any attempt, to afcertain the longitude of Yale-College; we made preparations, to avail ourselves of the opportunity furnished, by the eclipse of the sun, of Sept. 17th 1811. The going of the clock, from which the time was to be taken, was carefully attended to, for feveral months preceding. Its rate was found, by observing the passage of the sun and fixed stars, across the meridian. A transit instrument, with a telescope of about twenty inches focus, had been fixed, under the cupola of the Lyceum, nearly two years before. Care had been taken, during this period, by repeated observations of the pole star, to obtain an accurate adjustment of the telescope in the plane of the meridian.

When the meridian was fatisfactorily determined, an object was placed in view, nearly two miles diftant, by which the inftrument might afterwards be eafily adjus-

ted at pleasure.

The day of the eclipse was remarkably fine. Scarce a cloud was to be seen, the whole time the sun was above the horizon. There was little or no wind, to disturb our observations. One of us was stationed at the clock; while the other was looking at the sun, through a three seet resecting telescope. The eclipse was perceived, almost, if not quite, at the instant of its commencement. The time of the end was observed, with no less exactness.

The rate of the clock was determined, by observations on the meridian transits, of the sun, and of the fixed stars, on that and the preceding and following days. It was found to lose eight seconds, in twentyfour hours. The passage of the sun, across the meridian, on the day of the eclipse, was very carefully observed with the transit instrument. The time, by the clock, was 11h -54'-9.5"

The following were the observations of the beginning

and end of the eclipfe.

	h. m. s.
Beginning, by the clock,	0 38 22
Clock slower than the sun,	0 5 50.5
Allowance for the rate of the clock,	0 0 0.2
Apparent time of beginning,	0 44 12.7
Time of end, by the clock,	3 46 5
Clock slower than the sun,	5 50.5.
Allowance for the rate of the clock,	0 0 1.3
Apparent time of end,	3 51 56.8

To find the latitude of the place, the meridian altitude of the sun, was repeatedly taken, with an Equatorial Instrument, containing a telescope of 18 inches focus. The mean of twelve observations, was very nearly 41°—18'. The latitude is taken at this, in the following calculations; and the reductions of parallax

and latitude are made, according to that estimate of the figure of the earth, which gives the proportion of the polar to the equatorial diameter, as 300 to 301.

From these data, the longitude might be calculated, if the folar and lunar tables could be relied on, as perfectly correct. But it is well known that they are liable to an errour, which might materially affect the refult. It is necessary, therefore, that the tables should be corrected, or that the time of true conjunction should be obtained, from observations, at some place or places, whose longitude is already known. For this, we are indebted to a very obliging communication from Nathaniel Bowditch Efq. of Salem; on whose accuracy in calculation, full reliance may be placed: and who had undertaken to collect the observations, which were made on the eclipfe, in different parts of the United States. He finds the time of true conjunction to be at 6h. 57'. 05.8" apparent time at Greenwich; the longitude of the fun and moon, at that time, 173° 56' 32.4"; and the moon's Latitude 36' 40.2" North. lowing are the calculations for the longitude at Yale College.

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and the second of the second o	h. m. s.
Apparent time of beginning at Yale-College,	0 44 12.7
Assumed difference of Longitude,	4 51 50
Apparent time at Greenwich,	5 36 02.7
Sun's Right Ascension in time,	11 37 34
Sun past meridian,	0 44 12.7
Correction for the sun's advance,	+ 0.7
R. ascension of mid-heaven, in time,	12 21 47.4
Do in degrees, -	185° 26' 51"
Distance of meridian from Capricorn, -	84 33 09
Maria and Caraca and American State of the Caraca and C	h. m. s.
Apparent time of beginning	3 36 02.7
Do. of conjunction, (from Mr. Bowditch)	- 6 57 05.8
Time from beginning to conjunction, -	1 21 03.1
Moon's Lon. at conjunction,	173° 56′ 32″.4

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THE STREET STREET, STR	
Motion in 1 hour, (by Burg's tables,) - 0° 29' 32".; Do. in 21m. 3.1s 10 22.9	~ 520
Moon's Lon. at the beginning 173 16 36.8	100
Moon's Lat. at conjunction, 36 40.2	N
Motion in 1 hour 2 43.2 Do. in 21m. 3.1s 57.2	
Lat. at the beginning, 32 59.7	
Dist. merid. from Cap. 84 ⁶ 33' 09" Cos. 8.97742 Co. Lat. reduced, - 48 53 20 Tan. 10.05913	
Arc I 6 12 30 Tan. 9.03655 Dist. of the poles, - 23 27 42	63
Arc II 17 15 12 Sin.A.C. 0.52783 Arc I 6 12 30 Sin. 9.03400 Dist. of merid. from Cap. 84 33 09 Tan. 11.02061)22
Dist. of the nonagesimal from Cancer, Add 10.58244	192
Lon. of nonagesimal, 165 20 54 Lon. of the moon 173 16 36.8 Dist. of the moon E. fr. non. 7 56 02.8 Dist. of the non'l. fr. Cancer 75 20 34 Cos. 9.40318 Arc II. 17 15 12Tan.+10 19.49216	
Alt. of nonagesimal 50 49 44 Tan. 10.08898 Moon's equa. } 53' 59".9 = 3239".9 Reduc. for Lat. 41° 18' — 4.7	806
Moon's reduced hor. par. 3235.2 Sun's horizontal parallax - 8.7	
Hor. par. of moon from sun, 3226.5 Log. 3.50873 Altitude of the nonages. 50 49 44 Sin. 9.88944 Moon's Lat 0 32 59.7 Cos.A.C. 0.00003	491
Dist. of the moon from non. 7 56 02.8 Sin. 3.39820	1.6

10 TO 1 DOG 1 TO 1 DAY 10	17 C. L.			C. 19 C
43 20 20 20 20 20 20 20 20 20 20 20 20 20		Se -1	-	100
Longitude	-	Trail.	1000	A Section :
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Approximate par. in Lon.	5' 45.3"	Log.	2.5381876
Wall Vision Comments	and miles		3.3982008
Dist. of moon from non. +par. in lon. }	1 48.1	Sin.	9.1451717
Correct par. in lon. 349.5" =	5 49.5	Log.	2.5433725
Hor. par. of moon from sun, Alt. of the non 50	3226.5" 49 44	Log. Cos.	3.5087317 9.8004686
1st Part of par. in lat. 2038" = 2d Part	33 58 - 0.7		3,3092003
	33 58.7 32 59.7		Town to shell
Moon's apparent lat. Moon's hor. semid. 14' 44.2' Inflexion,	0 59	8.	se al la ar leg 180 à 180 à
Augment. for 50° alt. 4. 10.7		Alatadaga Ferrina Farrina	Total
Moon's cor. semid. 14 5	3 = 893"		o amin'ny tanàn
Sun's semid. 15 5 Irradiation, -	7.2 3.5		Bryan Carlon
Sun's cor. semid. 15	53.7 = 95	3.7	1.305040
Sun and moon's semidiameter Moon's appar. Lat.		6.7	Markette 48 Augustinasi
Sum, Difference,			3.280054 5 3.2522946
Appar. Lon. of moon from su Par. in lon.	in, 184	5.76 Lo	2)6.5323491 g. 3.2661745
True lon. of moon from sur	i, 21g	95.26	A STATE OF
Moon's hourly mo. in lon. 29 Sun's do.	2 32.7° 2 26.5 5	Andrews .	

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Hourly mo. fr. sun 27' 06.15 One hour,	"=1626.15" Log. A 3600 Log.	3.5563025
Lon. of moon from sun,	- 2195.26 Log.	3.3414860
Time from beginn. h. m. s to conjunction, 1 20 59	10-49r0 0 Tor	3.6866279
Time of begining, 0 44 15	3.7 Log.	3.0800279
Time of coni at V C 9 05 1	o c	Mach s late av

For the Parallaxes, &c. at the end.

Apparent time of end, Assumed differ. of lon.	h. m. s. 3 51 56.8 4 51 50	Moon's int. at the ra Dist aicrid, from an Do. lat. reduced,
Apparent time at Greenwi Sun's Right-ascen. in time	11 38 02	Arc. I. Dist. of the poles.
Sun past meridian, *Cor. for the sun's advance	3 51 56.8 e, 4 3.4	Arra IL.
R. Ascen. of mid heaven	15 30 02.2	Dist of merid, from
Do. in degrees, Dist. of merid. from cap.	232° 30° 33″ 37 29 27 b. m. s.	Disk non From cap. Subtract this from
Appa. time of end,	8 43 46.8	Lon magges.

This correction, which depends on the difference in the equation of time, on two fuccessive days, is generally emitted, in the rules which are given, for finding the right ascension of midheaven; probably because it is of small account. But, in strict propriety, it ought to be included. In the present instance. it will make a difference, of more than a minute, in the longitude of the nonagefimal. The equation of time is 21 feconds greater, on the 18th of Sept, than on the 17th. The fun therefore, in twenty four hours, not only makes a complete revolution; but, in addition to this, falls to the west of the meridian, such a distance, as corresponds to 21 seconds of time. A proportional part of this, should be allowed, for the time between 12 o'clock, and the end of the eclipse.

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Do. of conjunction,	Lacal 6 57	05.8	Housig san fr
Time from conjunc. to Moon's lon. at conjunc. Motion in one hour,	173° 56	5' 32.4"	The hour, Loss of moon! Their from her
Do. in 46m. 41s.		22 59.6	office conjunction
Moon's long, at the end Moon's lat. at conjunc, Motion in one hour. Do. in 46m. 41s.	0.83	49 05.2	Tipe of conj. a
Moon's lat. at the end,	and the second	1 30	1- 15-14-113
Dist. merid. from cap.		" Cos.	9.8995200
Co. lat. reduced,	48 53 20	Tan,	10.0591364
Arc. I.	42 16 37	Tan.	9.9586564
Dist. of the poles,	23 27 42	g ein 4, .50 c	nest of the
Arc, II,	18 48 55	Sin. A	C. 0.4914459
Arc. I.	42 16 37	-	9.8278319
Dist. of merid. from cap			9.8848366
Dist. non. from cap. Subtract this from	57 59 39 270		10.2041135
Lon. nonages.	212 0 2		to sum agg
Lon. of the moon,	174 49 0		
th depends on the	dia man	lus correct	Nore.
Moon from non.	37 11 1		Mercage in t
Non. from cap. Arc. II.	57 59 39 18 48 55		9.724 2805
u odpeaven; probe	10 40 00	s mg en	0.13,3021011
Alt. of non-	32 44 09	Tan.	9.8081239
Hor. par. of moon fr. su	n. 3227	.2" Log.	3.5088124
Alt. of nonages.	32 44 09		9.7330098
Moon's lat.	0 41 30	Cos, A.	C. 0.0000316
sense tool ylang	ni panjar	Log.	3.2418538
Moon from non,	37 11 16		9.7813454
Approx. par. in long.	17 35	Log.	3,0231992
som sal for her ringe	osls od blu	f. this, the	may anner
die eglinte.			
Mo. from non. 7 par.	37 28 51	i Sin.	9.7842577

Hor. par. mo. from sun	17"42" Log. 3227.2 Log. 2 44 09 Cos.	3.0261115 3.5088124 9.9248852
1st part of par. in lat. 2d part of do.	45 14.54 Log.	3.4336976
Cor. par. in lat. Moon's true lat.	45 14.7 41 30	Dif of L
Moon's appar. lat. 224.7 M's. semidinflex. 14 42.5 Aug. for 25° 30′ alt. 46.1	and the second s	in an arthur
Mo's. cor. semid. 14 48.6 Sun's do. 15 53.		and plants of
Sun and mo's. sem. 30 42. Mo's. app. lat.	3 = 1842.3" 224.7	to desired
Sum Difference,	2067. Log. 1617.6 Log.	3.3153403 3.2088711
		2)6.5242116
Appar. long. of moon from sun Par. in Lon.	1828.55 Log.	44
Far. in Lon.	1062.	
True long. of moon }	2890.55	
Mo's. hourly mot. 29 33.1. Sun's do. 2 26.5	2	
M's. hourly mot. \27 06.	6 = 1626.6 Log.A	.C. 6.788719
One hour,	3600 Log.	3.556302
Long. of m. fr. sun	2890.55 Log.	3.460980
Time fr. conj. to end 1 46 App. time of end 3 51		. 3.806002

App. time of conj. Do. cal. from begin.	2 05 1 2 05 1	9.4	T Went	on long	4 2 *
Mean,	2 05		P.	der son	to IIA
Conj. at Greenwich	6 57	05.8	Service .	or sue to	rido nig
Mer de der	h	m	200	0 413 4	10

Dif. of long. 4 51 49.8=72 57 27.

It would have been defirable, to obtain a correction of the tables, from observations, made at the first meridian at Greenwich; that our calculation might not be affected, by the small errours, to which the longitudes of all places in the United States are liable. But as this eclipse was not visible in Europe, we have had recourse to observations made here, to determine the latitude and longitude of the moon, and the true time of conjunction.

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Yale College, Nov. 24, 1812.

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